



Infrared spectroscopy of radioactive hydrogen chloride H³⁶Cl

Santeri Larnimaa^{a,*}, Markku Vainio^{a,b}, Ville Ulvila^c

^a Department of Chemistry, University of Helsinki, P.O. Box 55, Helsinki 00014, Finland

^b Photonics Laboratory, Physics Unit, Tampere University, Tampere, Finland

^c VTT Technical Research Centre of Finland Limited, Espoo, Finland



ARTICLE INFO

Article history:

Received 25 August 2021

Revised 18 October 2021

Accepted 19 October 2021

Available online 29 October 2021

Keywords:

Radioactive hydrogen chloride H³⁶Cl
Fourier-transform infrared spectroscopy
FTIR
Rotational-vibrational spectroscopy
Line center wavenumber
Molecular constant

ABSTRACT

We present the first report of optical absorption spectroscopy of H³⁶Cl, a radioactive isotopologue of hydrogen chloride. We used Fourier-transform infrared spectroscopy to determine the line center wavenumbers of the fundamental rovibrational band lines P(10)–R(10) and the first overtone band lines P(1)–R(7) with total uncertainty of less than 0.0018 cm⁻¹ (60 MHz) and 0.007 cm⁻¹ (0.2 GHz), respectively, at 68% confidence level. We also performed a rotational analysis on the bands to determine the related molecular constants. We further compared the linewidths and relative intensities of the lines to those of the stable isotopologues H³⁵Cl and H³⁷Cl. The new spectroscopic information assists in developing optical instrumentation for the detection of H³⁶Cl.

© 2021 The Authors. Published by Elsevier Ltd.

This is an open access article under the CC BY license (<http://creativecommons.org/licenses/by/4.0/>)

1. Introduction

Hydrogen chloride is an important molecule that plays a role in atmospheric chemistry [1,2], astrochemistry [3–5], the semiconductor industry [6], and in the research of volcanic activity [7–10] and planetary atmospheres [11–13]. While the rovibrational bands of HCl were studied in detail already in the 1960s [14–16], early works on HCl date to the beginning of the 20th century to the development of early spectrometers and the acceptance of quantum mechanics [17–19]. In recent years, there have been many publications that report increasingly accurate line parameters [20–24]. The naturally occurring stable isotopologues H³⁵Cl, H³⁷Cl, and their deuterated forms are usually studied. However, to the best of our knowledge, there are no prior reports on optical spectroscopy of H³⁶Cl.

Chlorine-36 is a radioactive isotope of chlorine that is formed naturally in the atmosphere [25]. It exists in miniscule concentrations in groundwater (background ³⁶Cl/Cl ratio is in the order of 10⁻¹⁵) but can be used for groundwater dating [26–29]. The species is also formed in nuclear facilities by neutron activation of ³⁵Cl impurities found in reactor materials [30–33]. This is a major concern due to the long (approx. 301,000 years) half-life of ³⁶Cl. Some of ³⁶Cl may be released to the environment in the form of H³⁶Cl.

Owing to their high sensitivity and selectivity, optical spectroscopy methods are well suited for the detection of radioactive compounds in the gas phase [34]. These methods can also offer a more affordable and portable alternative to the current state-of-the-art method, Accelerator Mass Spectrometry [35]. In addition, spectroscopic methods do not depend on the specific activity of the species of interest, which might lead to less laborious sample preparation than in Liquid Scintillation Counting [36], as the need for a complex radiochemical separation is circumvented.

Prior research on optical spectroscopy of radioactive compounds has focused especially on laser spectroscopy of radiocarbon dioxide ¹⁴CO₂ [34,37–45]. One goal has been to develop highly sensitive and portable field instruments for emission monitoring of this species in nuclear facilities and their decommission sites or for discriminating between emissions from the combustion of fossil and biofuels. In general, the development of laser spectroscopy instrumentation for the detection of radioactive molecules is hindered by the lack of spectroscopic information. For example, we only recently reported the first absorption measurements and rotational analysis of radiocarbon methane ¹⁴CH₄ [46,47], another important species of concern in nuclear facility emission monitoring [48].

Here we report the results of our Fourier-transform infrared spectroscopy (FTIR) measurements of the fundamental and first overtone rovibrational bands of H³⁶Cl. These results aid the development of optical instrumentation for monitoring this species. In the following sections, we first describe the sample preparation, measurement and line-fitting procedures, after which we summarize the main results. Finally, we present the rotational analysis

* Corresponding author.

E-mail address: santeri.larnimaa@helsinki.fi (S. Larnimaa).

and discuss the results for the overtone band. Supplementary information discusses the experimentally observed linewidths and relative intensities of the H^{36}Cl fundamental band lines, including a comparison to those of H^{35}Cl .

2. Sample preparation and measurement conditions

The gaseous H^{36}Cl sample was prepared by pipetting 50 μl of ^{36}Cl -enriched NaCl solution (3.7 MBq/ml activity concentration; American Radiolabeled Chemicals, Inc.) into a 10 cm long IR quartz cell (Suprasil 300; FireflySci Type 34). The aqueous sample was subsequently dried to solid NaCl using a flow of compressed air, after which a few drops of concentrated sulfuric acid were added and the cell was closed with PTFE stoppers. Finally, the reaction of the solid NaCl with sulfuric acid to produce HCl was initiated by tilting the sample cell to combine the reagents. This simple gas cell system was not equipped with a pressure gauge or a temperature meter to minimize the sample cell volume and sample losses.

Given the specific activity of ^{36}Cl (1220 Bq/ μg), the total amount of ^{36}Cl in the sample was expected to be 4.2 μmol . Based on the FTIR measurements, we obtained 1.6 mbar of gaseous H^{36}Cl . This estimate assumes the HITRAN H^{35}Cl line intensities for H^{36}Cl (see Supplementary Note 4). This means that the yield into gas phase was 43%, assuming 296 K temperature and 28 ml volume of the sample cell. The sample also contained approximately 0.4 mbar and 0.8 mbar of the stable isotopologues, H^{35}Cl and H^{37}Cl , respectively. The total partial pressure of HCl was 2.7 mbar. Since the pressure increase in the sample cell due to the produced HCl is small and the partial pressure of water in concentrated sulfuric acid is negligible [49], it is reasonable to assume 1 atm (1013.25 mbar) pressure in the cell during the measurements. The lab pressure was within a few mbar of this pressure upon filling the sample cell.

3. Measurement of the fundamental band

The H^{36}Cl measurements in the mid-infrared (MIR) were executed using a commercial FTIR instrument (Bruker IFS 120 HR). We used a Globar light source, Ge-on-KBr beam splitter, and liquid-nitrogen-cooled InSb detector. The raw spectrum is a result of processing 10 co-added double-sided interferograms using the Mertz method [50]. We used the manufacturer-specified resolution setting of 0.02 cm^{-1} (600 MHz). The manufacturer defines the resolution as the full width at half maximum (FWHM) of the assumed instrument line shape (ILS) function if triangular apodization was used; we used Norton-Beer medium apodization [51], which corresponds to 0.0095 cm^{-1} (285 MHz) half width at half maximum (HWHM) of the resulting ILS function. Table 1 lists other relevant FTIR instrument settings used.

Table 1
Relevant FTIR instrument settings used for the MIR measurements.

Light source	Globar
Beam splitter	Ge-on-KBr
Detector	InSb
Aperture diameter	1 mm
Apodization function	Norton-Beer medium
Assumed HWHM of ILS function	0.0095 cm^{-1} (285 MHz)
Number of co-added interferograms	10
Type of interferograms	Double-sided
Interferogram processing method	Mertz
Zero-filling factor	2
Optical band-pass filtering	2614–3378 cm^{-1}
Digital band-pass filtering	10–20 kHz (1975–3950 cm^{-1})
HeNe reference laser down-converted frequency	80 kHz

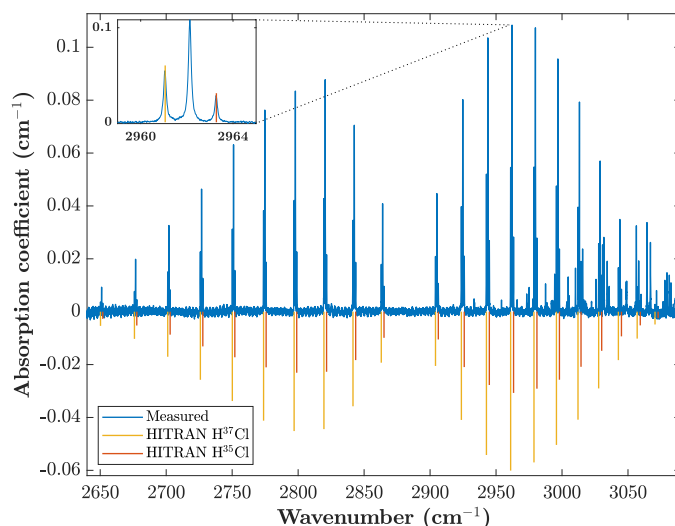


Fig. 1. The measured HCl absorption spectrum (blue). The middle peak of each of the peak triplets belongs to H^{36}Cl as revealed by the included H^{37}Cl (yellow) and H^{35}Cl (orange) HITRAN lines (mirrored for clarity). The inset shows a zoomed view of the R(3) lines. In addition to the HCl peaks, there are interfering air absorption lines (especially in the R-branch), and an etalon effect caused by the sample cell windows. Note that the amounts of the stable isotopologues H^{37}Cl and H^{35}Cl in the sample do not follow their natural abundances. The H^{35}Cl lines are thus the weakest in the spectrum.

The 100% transmission baseline was determined manually from the HCl spectrum by first low-pass filtering the spectrum, excluding the HCl peaks from the data, and finally by using Savitzky-Golay filtering [52] and defining the result as the baseline. The spectrum was then divided by this baseline to obtain the HCl transmission spectrum. The result was further converted into absorption spectrum using the Napierian Beer-Lambert law [53]

$$\alpha(\tilde{\nu}) = -\frac{1}{L} \ln \left(\frac{I(\tilde{\nu})}{I_0(\tilde{\nu})} \right),$$

where $\alpha(\tilde{\nu})$ is the absorption coefficient, L is the absorption path length (10 cm), $I(\tilde{\nu})$ is the raw spectrum, and $I_0(\tilde{\nu})$ is the 100% transmission baseline. The resulting absorption spectrum is shown in Fig. 1.

The absorption spectrum in Fig. 1 shows a peak structure characteristic to the fundamental rovibrational band of HCl. Instead of the typical H^{37}Cl and H^{35}Cl peak doublets, we see triplets whose middle peak of each belongs to H^{36}Cl . We only consider lines P(10)–R(10) in this article, as lines further from the band center were masked by noise. The strongest H^{36}Cl peak is line R(3) with 0.108 cm^{-1} peak absorption (34% in transmission) and the weakest is line R(10) with 0.008 cm^{-1} peak absorption (92% in transmission). In addition to the strong HCl lines, attention in Fig. 1 is drawn to the interfering air absorption lines and to the etalon effect. The relatively strong air absorption lines are explained by the fact that the FTIR instrument was not evacuated or purged. The free spectral range of the etalon is typically 2.3–2.8 cm^{-1} (70–80 GHz), which matches well with the 1.25 mm thickness of the IR quartz sample cell windows. The etalon and the interfering air absorption lines are included in the fit model as discussed below.

4. Fitting procedure

Due to the relatively high measurement pressure (1 atm), pressure broadening dominates most of the transition linewidths (see Supplementary Note 1). For this reason, we chose to model the absorption peaks as Lorentzian functions. All of the 21 rotational components (lines P(10)–R(10)) of the H^{36}Cl fundamental band

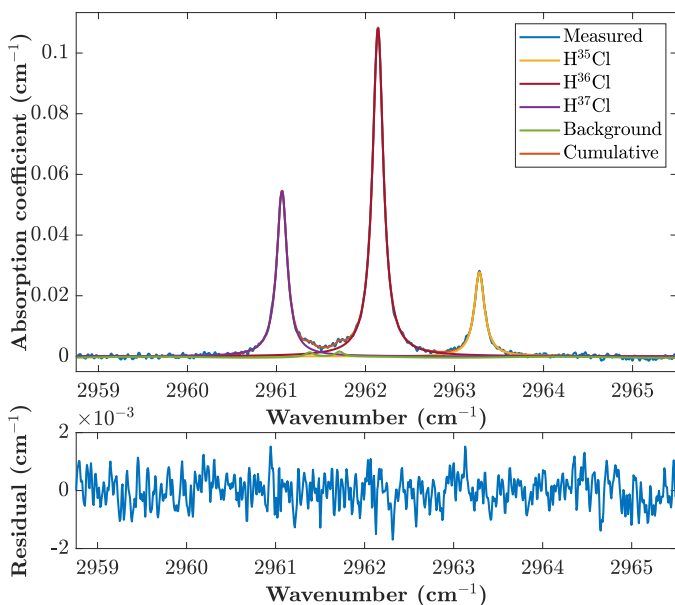


Fig. 2. Fit to the measured R(3) lines (upper panel) and the fit residual (lower panel). The three major peaks belong to H³⁷Cl (leftmost), H³⁶Cl (middle), and H³⁵Cl (rightmost). The background fit consists of a constant, a sine function, and two Lorentzian functions for the two small air absorption peaks at approximately 2961.4 cm⁻¹ and 2961.7 cm⁻¹.

were fitted separately such that a typical fitting window was 4–9 cm⁻¹ around the H³⁶Cl peak in question. To give an example, Fig. 2 shows the fit for line R(3). The fit model consists of three Lorentzian functions (one for each HCl isotopologue), a constant for a possible residual background offset, and a sine function to consider the etalon effect. In addition, extra Lorentzian functions were added for air absorption lines when necessary; in the case of the R(3) line, two small air peaks were fitted (Fig. 2).

The lack of systematic features in the residual in Fig. 2 shows that the chosen fit model indeed describes the data well. The signal-to-noise ratio of the H³⁶Cl peak is 212, defined as the ratio of the maximum peak absorption to the standard deviation of the fit residual. The peaks with the lowest SNR that were included in the analysis of the fundamental band are P(10) and R(10) (SNR≈15). Note that we also calibrated the wavenumber axis using the HITRAN zero pressure wavenumbers and the related 1 atm pressure shifts of the H³⁵Cl and H³⁷Cl P(10)–R(10) lines [20,54,55]. The analysis revealed an average offset of -0.0101 cm⁻¹ (-303 MHz) between the fitted H³⁵Cl and H³⁷Cl line center wavenumbers and the corresponding HITRAN values. Note that the average offset was calculated by weighting the data points by the inverses of the corresponding fit uncertainties. A similarly weighted standard deviation of the offsets is 0.00124 cm⁻¹ (37 MHz), which was combined with the fit uncertainties of the line center wavenumbers (see Results).

5. Results

Table 2 lists the fit results for the H³⁶Cl line center wavenumbers. The numbers in parentheses in Table 2 are one standard deviation uncertainties in least significant digits as obtained from the fit. The total uncertainty column contains the combined effect of the fit uncertainties and the calibration uncertainty at 68% confidence level. Other uncertainty sources, such as the uncertainty of the baseline determination or the measurement pressure via the pressure shifts, are negligible. Overall, the total wavenumber uncertainties are below 0.0018 cm⁻¹ (60 MHz).

Table 2

Line center wavenumber $\tilde{\nu}$ results for the H³⁶Cl fundamental band. The numbers in parentheses are one standard deviation fit uncertainties in least significant digits and the values in the column denoted by $\sigma \tilde{\nu}$ represent the combined effect of the fit and calibration uncertainties at 68% confidence level. The symbol $\tilde{\nu}_{\text{calc}}$ refers to the calculated line center wavenumbers as obtained from the rotational analysis discussed in section 6. Note that all the values in the table are given at 1 atm pressure and 296 K temperature.

Line	$\tilde{\nu}$ (cm ⁻¹)	$\sigma \tilde{\nu}$ (10 ⁻⁴ cm ⁻¹)	$\tilde{\nu} - \tilde{\nu}_{\text{calc}}$ (10 ⁻⁴ cm ⁻¹)
P(10)	2651.0675(13)	18	9
P(9)	2676.8101(6)	14	0
P(8)	2702.0639(4)	13	-6
P(7)	2726.8168(3)	13	3
P(6)	2751.0527(2)	13	10
P(5)	2774.7574(2)	13	-1
P(4)	2797.9207(2)	13	0
P(3)	2820.5297(2)	13	-5
P(2)	2842.5712(3)	13	7
P(1)	2864.0247(4)	13	1
R(0)	2905.1493(4)	13	-1
R(1)	2924.7790(2)	13	-2
R(2)	2943.7809(2)	13	-4
R(3)	2962.1397(2)	13	6
R(4)	2979.8392(2)	13	1
R(5)	2996.8709(2)	13	-1
R(6)	3013.2251(2)	13	7
R(7)	3028.8887(3)	13	4
R(8)	3043.8510(4)	13	3
R(9)	3058.1013(6)	14	2
R(10)	3071.6279(13)	18	-9

While determination of the line center wavenumbers was the main objective of this research, the measurements also provided information about the widths and relative intensities of the lines. These results are discussed separately in Supplementary information. Briefly, the relative intensities of the H³⁶Cl seem to follow those of the stable isotopologues H³⁵Cl and H³⁷Cl as expected based on previous research on the stable isotopologues. However, our results for the pressure-broadened linewidths are systematically larger than the HITRAN reference data for the stable isotopologues.

6. Rotational analysis

We retrieved the molecular constants for the H³⁶Cl fundamental rovibrational band. The $\nu = 0$ band line center wavenumbers of a diatomic molecule can be approximately expressed as [56]

$$\begin{aligned} \tilde{\nu} = & \tilde{\nu}_{\nu-0} + (\tilde{B}_{\nu} + \tilde{B}_0)m + (\tilde{B}_{\nu} - \tilde{B}_0 - \tilde{D}_{\nu} + \tilde{D}_0)m^2 \\ & + (\tilde{H}_{\nu} + \tilde{H}_0 - 2\tilde{D}_{\nu} - 2\tilde{D}_0)m^3 \\ & + (3\tilde{H}_{\nu} - 3\tilde{H}_0 - \tilde{D}_{\nu} + \tilde{D}_0)m^4 \\ & + (3\tilde{H}_{\nu} + 3\tilde{H}_0)m^5 + (\tilde{H}_{\nu} - \tilde{H}_0)m^6, \end{aligned} \quad (1)$$

where $\tilde{\nu}_{\nu-0}$ is the band center, \tilde{B} is the rotational constant, and \tilde{D} and \tilde{H} are the different order centrifugal distortion constants. The integer m is equal to $-J$ for the P-branch and $J + 1$ for the R-branch. The rotational quantum number J refers to the lower state.

We could not determine the centrifugal distortion constants \tilde{H} in a statistically significant manner; we thus assumed $\tilde{H} = \tilde{H}_{\nu} = \tilde{H}_0$ and constrained its value to 1.66742×10^{-8} cm⁻¹. This value is the predicted \tilde{H}_0 of H³⁶Cl, which we calculated using \tilde{H}_0 , \tilde{H}_1 , and \tilde{H}_2 of H³⁵Cl given in the supplementary material of Ref. [20], the approximate Dunham expression [57,58]

$$\tilde{C}_{\nu,j} = \tilde{Y}_{0j} + \tilde{Y}_{1j}(j + \frac{1}{2}) + \tilde{Y}_{2j}(j + \frac{1}{2})^2 \quad (2)$$

and the approximate isotope relation [15,57]

$$\frac{\tilde{Y}_{ij}^s}{\tilde{Y}_{ij}} = \rho^{i+2j}, \quad \rho = \sqrt{\frac{\mu}{\mu^s}}, \quad (3)$$

Table 3

The experimental and predicted fundamental band molecular constants for H³⁶Cl obtained as explained in the text. The values refer to zero pressure.

Parameter	Experimental (cm ⁻¹)	Predicted (cm ⁻¹)
$\tilde{\nu}_{1-0}$	2884.8936(3)	2884.8933
\tilde{B}_0	10.43210(4)	10.43213
\tilde{B}_1	10.12843(4)	10.12847
$\tilde{D}_0/10^{-4}$	5.275(3)	5.273
$\tilde{D}_1/10^{-4}$	5.207(3)	5.206
$\tilde{H}_0/10^{-8}$		1.66742

where \tilde{Y}_{ij} are Dunham coefficients [57], μ is the reduced mass, and $\tilde{C}_{\nu,0} - \tilde{C}_{0,0} = \tilde{\nu}_{\nu-0}$, $\tilde{C}_{\nu,1} = \tilde{B}_\nu$, $\tilde{C}_{\nu,2} = \tilde{D}_\nu$, and $\tilde{C}_{\nu,3} = \tilde{H}_\nu$. The constants marked with an asterisk (*) refer to the heavier isotope.

Eqs. (2) and (3) and the molecular constants of H³⁵Cl given in the supplementary material of Ref. [20] were further used to predict in a similar manner the remaining molecular constants for H³⁶Cl. The results are listed in Table 3 together with the experimental molecular constants that we determined from the measured line center wavenumbers of lines P(10)–R(10) in least-squares fits using the combination relations [15,56]

$$R(J-1) - P(J+1) = (4\tilde{B}_0 - 6\tilde{D}_0 + \frac{27}{4}\tilde{H})\left(J + \frac{1}{2}\right) - (8\tilde{D}_0 + 34\tilde{H})\left(J + \frac{1}{2}\right)^3 + 12\tilde{H}\left(J + \frac{1}{2}\right)^5 \quad (4)$$

$$R(J-1) + P(J) = 2\tilde{\nu}_{1-0} + 2(\tilde{B}_1 - \tilde{B}_0 - (\tilde{D}_1 - \tilde{D}_0))J^2 - 2(\tilde{D}_1 - \tilde{D}_0)J^4, \quad (5)$$

where $J = 1, \dots, 9$ for Eq. (4) and $J = 1, \dots, 10$ for Eq. (5).

First, the lower-state parameters \tilde{B}_0 and \tilde{D}_0 were determined in a least-squares fit using Eq. (4) and the assumption $\tilde{H} = 1.66742 \times 10^{-8} \text{ cm}^{-1}$. However, prior to fitting, we shifted the measured H³⁶Cl line center wavenumbers to those corresponding to zero pressure assuming the HITRAN H³⁵Cl pressure shifts [54,55]. After this, we determined the band center $\tilde{\nu}_{1-0}$ and upper-state parameters \tilde{B}_1 and \tilde{D}_1 using Eq. (5). The lower-state parameters were fixed during the least-squares fit to those obtained from the fit of Eq. (4). In addition, to consider the fit uncertainties of the lower-state parameters in the determination of the upper-state parameters, we repeated the fit a hundred times, each time drawing new \tilde{B}_0 and \tilde{D}_0 values from normal distributions whose expectation values and standard deviations were chosen to be the original \tilde{B}_0 and \tilde{D}_0 values and their respective fit uncertainties obtained from the fit of Eq. (4). Finally, we calculated the standard deviations of the hundred new \tilde{B}_0 and \tilde{D}_0 values and combined them with the fit uncertainties obtained from the original fit of Eq. (5) to estimate the total uncertainties of the upper-state parameters.

The experimental and predicted molecular constants in Table 3 are in excellent agreement. This is also true for the observed minus calculated line center wavenumber values in column 4 of Table 2 (denoted by $\tilde{\nu} - \tilde{\nu}_{\text{calc}}$). These values were calculated using Eq. (1), the experimental constants in Table 3, and the assumption $\tilde{H} = \tilde{H}_\nu = \tilde{H}_0$; the pressure shifts were reapplied before the comparison. The largest difference is 0.0010 cm^{-1} (30 MHz) and all the values are within the estimated total uncertainties.

Although the experimental and calculated wavenumbers are in excellent agreement, it is well known that the molecular parameters obtained with this traditional approach may perform poorly in predicting accurate line center wavenumbers for transitions with high quantum number J values [20]. Indeed, we repeated the above-explained procedure to determine the molecular constants also for H³⁵Cl and H³⁷Cl. Even though the result parameters were within two standard deviations from those obtained with more refined methods [20] or from highly accurate experiments con-

sisting of a considerably larger dataset [59], our model for H³⁵Cl (H³⁷Cl) predicts up to 0.050 cm^{-1} or 1.5 GHz larger (0.020 cm^{-1} or 0.6 GHz lower) wavenumber values for lines P(15) and R(15) when compared to the corresponding HITRAN line center wavenumbers. We can then expect similar inaccuracy for the H³⁶Cl model.

7. The first overtone band

We also measured the first overtone 2 – 0 rovibrational band of H³⁶Cl. Due to the limited amount of Na³⁶Cl solution available for us at the time, we conducted the measurements with the same sample that we used in the MIR measurements. For these near infrared (NIR) measurements, we co-added 100 interferograms and used a Si-on-CaF₂ beam splitter, Tungsten light source, and a liquid-nitrogen-cooled InSb detector. The measurement settings, baseline determination, and fitting procedure were similar to what we used in the MIR measurements. The relevant measurement settings are listed in Table 4.

Despite the longer averaging time, the SNR in the NIR measurements is worse than in the MIR measurements. This was expected, since the HCl overtone transitions are typically 30–50 times weaker than the corresponding fundamental ones. The P-branch is also considerably interfered by air absorption. For these reasons, we could only reliably analyze lines P(1)–R(7), and only the H³⁷Cl lines R(1)–R(7) were used in the wavenumber axis calibration. The calibration revealed an average offset of -0.01479 cm^{-1} (-443 MHz) compared to the corresponding HITRAN values [20,54]. However, because the HITRAN pressure-shift data is extrapolated from the MIR measurements by Pine and Looney [55], we chose to use the NIR pressure-shift data by Asfin et al. [23] to calculate the H³⁷Cl line center wavenumbers at 1 atm. The standard deviation of the offsets is 0.00364 cm^{-1} (109 MHz) and was again interpreted as the calibration uncertainty. Note that in the calculation of the average and the standard deviation of the offsets, the data points were again weighted using the inverses of the fit uncertainties. As an example, Fig. 3 shows the overtone line R(3) fit. This is the strongest H³⁶Cl line whose peak absorption is 0.0027 cm^{-1} (97.3% in transmission) with an SNR of 20; the lowest SNR is 7 and for line P(1) with 0.0010 cm^{-1} peak absorption (99.0% in transmission).

Table 5 lists the fit parameter results for the H³⁶Cl first overtone band. Due to the generally modest SNR, we could not determine the linewidths and relative intensities reliably, for which reason we report only the line center wavenumbers. The numbers in parentheses are the one standard deviation fit uncertainties in least significant digits. The total uncertainty column values represent the combined fit and calibration uncertainties at 68% confidence level. They are less than 0.007 cm^{-1} (0.2 GHz). The observed minus calculated column values in Table 5 are again in good

Table 4

Relevant FTIR instrument settings for the NIR measurements.

Light source	Tungsten
Beam splitter	Si-on-CaF ₂
Detector	InSb
Aperture diameter	1 mm
Apodization function	Norton-Beer medium
HWHM of ILS function	0.0095 cm ⁻¹ (285 MHz)
Number of co-added interferograms	100
Type of interferograms	Double-sided
Interferogram processing method	Mertz
Zero-filling factor	2
Optical band-pass filtering	6667–5000 cm ⁻¹
Digital band-pass filtering	25–35 kHz (4937–6912 cm ⁻¹)
HeNe reference laser down-converted frequency	80 kHz

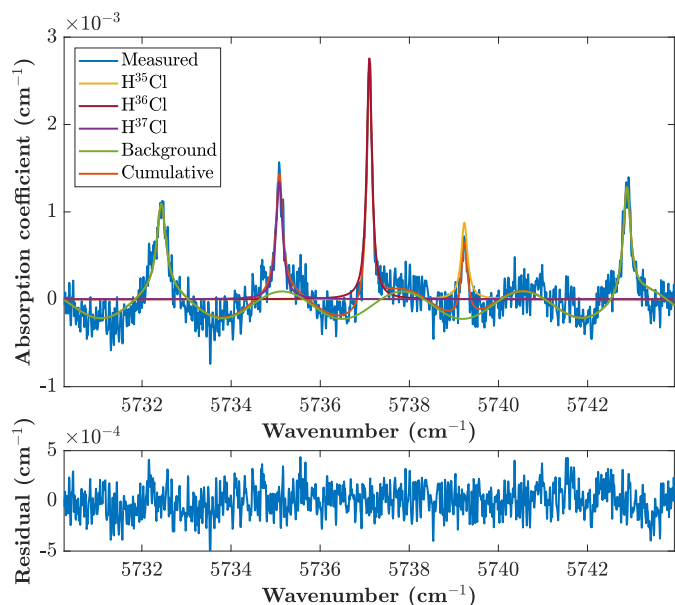


Fig. 3. Fit to the measured first overtone R(3) lines (upper panel) and the fit residual (lower panel). The leftmost and rightmost peaks are fitted air absorption lines and the remaining three fitted peaks belong to H³⁷Cl, H³⁶Cl, and H³⁵Cl in the order of increasing wavenumber. In addition to the two air absorption lines, the background function consists of a constant and a sine.

Table 5

Line parameter results for the H³⁶Cl first overtone band. The line center wavenumber is denoted by $\tilde{\nu}$. The numbers in parentheses are one standard deviation fit uncertainties in least significant digits and the values in the column denoted by $\sigma\tilde{\nu}$ represent the combined effect of the fit and calibration uncertainties at 68% confidence level. The symbol $\tilde{\nu}_{calc}$ refers to the calculated line center wavenumbers as obtained from the rotational analysis discussed in the text. Note that all the wavenumber values in the table are given at 1 atm pressure.

Line	$\tilde{\nu}$ (cm ⁻¹)	$\sigma\tilde{\nu}$ (10 ⁻³ cm ⁻¹)	$\tilde{\nu} - \tilde{\nu}_{calc}$ (10 ⁻³ cm ⁻¹)
P(1)	5645.043(5)	7	21
R(0)	5685.525(5)	6	-21
R(1)	5703.965(3)	5	-5
R(2)	5721.168(2)	5	3
R(3)	5737.113(2)	4	-1
R(4)	5751.802(2)	4	-1
R(5)	5765.227(2)	4	2
R(6)	5777.370(3)	5	3
R(7)	5788.217(4)	5	-2

Table 6

Experimental and predicted first overtone band molecular constants for H³⁶Cl obtained as explained in the text. The values refer to zero pressure.

Parameter	Experimental (cm ⁻¹)	Predicted (cm ⁻¹)
$\tilde{\nu}_{2-0}$	5665.895(10)	5665.898
\tilde{B}_2	9.8274(8)	9.8272
$\tilde{D}_2/10^{-4}$	5.18(11)	5.15

agreement with the experiment and deviate less than 0.005 cm⁻¹ (150 MHz) from the experimental values, not including the low SNR lines P(1) and R(0), for which large deviations are apparent. The observed minus calculated values have been obtained from a similar rotational analysis to what was discussed for the MIR measurements. However, due to the small number of P-branch lines, we used Eq. (1) directly in the least-squares fit, made the $\tilde{H}_v = \tilde{H}_0 = 1.66742 \times 10^{-8}$ cm⁻¹ assumption, and fixed the remaining lower-state parameters to the experimental values in Table 3.

Table 6 lists the experimental and predicted molecular constants obtained for the first overtone band in a similar manner as discussed for the fundamental band. The accuracy of the experimental parameters is not sufficient to ensure accurate line center wavenumber predictions for high J values. We advise using the predicted molecular constants instead. For example, similarly predicted molecular constants for H³⁷Cl predict up to 0.027 cm⁻¹ (0.8 GHz) larger line center wavenumbers for lines P(15) and R(15) compared to HITRAN. We note that the predicted molecular constants for H³⁷Cl agree within at least five significant decimals compared to the literature values [20] not including the band center, which agrees within 0.0011 cm⁻¹ (33 MHz).

8. Conclusions and outlook

In this article, we reported the first absorption spectrum of H³⁶Cl. The measurements of the fundamental and the first overtone rovibrational bands were performed using a commercial FTIR instrument with a thermal light source and using a sample prepared from the reaction between sulfuric acid and ³⁶Cl-enriched NaCl salt. The simple gas cell system helped to maximize the number density of the species of interest inside the cell, which provided us with strong absorption lines in the MIR and allowed us to determine the fundamental band P(10)–R(10) line center wavenumbers (referenced to the HITRAN values of the stable isotopologues H³⁵Cl and H³⁷Cl), with total uncertainties of less than 0.0018 cm⁻¹ (60 MHz) at 68% confidence level. Due to the order of magnitude weaker molecular transitions and consequently modest SNR in the NIR, we could determine the first overtone band line center wavenumbers of only lines P(1)–R(7) and with total uncertainties of less than 0.007 cm⁻¹ (0.2 GHz) at 68% confidence level.

The simple gas cell system came with the downside of not knowing the exact pressure and temperature inside the sample cell. However, the initiation of the reaction at atmospheric pressure and room temperature made it possible to estimate the measurement conditions with reasonable accuracy. In the Supplementary information, we investigated the linewidths and relative intensities of the H³⁶Cl lines and compared them to the corresponding literature values of the stable isotopologues H³⁵Cl and H³⁷Cl. The relative intensities showed similar behavior as the reference HITRAN data, implying that the line intensities of the different isotopologues are similar as expected based on previous research on the stable isotopologues. However, as discussed in the Supplementary information, the linewidths for all three isotopologues were systematically larger than the HITRAN reference data for H³⁵Cl and H³⁷Cl.

An advantage of having atmospheric pressure inside the sample cell during measurements was to minimize the distorting effect of the ILS function and to allow us to use a simple Lorentzian fit model. However, precise control of the measurement pressure and measurements in lower pressures to reduce the linewidths would enable much more accurate line center wavenumber determinations. To increase measurement sensitivity, an external coherent light source in the FTIR instrument instead of the thermal source could be useful, especially if combined with photoacoustic detection [47]. For accurate linewidth determinations, eliminating the ILS function altogether would be helpful. This could be done by using optical frequency combs [60] or by avoiding FTIR altogether (i.e., by using tunable laser absorption spectroscopy instead) [55,61,62]. In particular, the sensitivity improvement by laser absorption spectroscopy would allow more accurate characterization of the first overtone band. In addition, determination of the absolute line intensities would require a more elaborate sample preparation and gas-control system to ensure accurate external knowledge of the sample concentration [22]. Finally, as future applications in monitoring H³⁶Cl will expectedly require very

high sensitivities to detect trace amounts of this species, it would be valuable to demonstrate low-concentration measurements using some highly sensitive laser spectroscopy method, such as cavity ring-down spectroscopy [34,39] or photoacoustic spectroscopy [37,46,63]. These pursuits are assisted by our results, as we now have reasonably accurate knowledge on the line center wavenumbers of H³⁶Cl, and the simple models from the rotational analyzes can be used to predict line positions of yet unobserved transitions.

Funding

This work obtained funding from the Euratom research and training programme 2014–2018 [Grant agreement No. 755371, the CHANCE project]; the Academy of Finland [Project No. 326444]; and the Flagship on Photonics Research and Innovation of the Academy of Finland (PREIN). S. Larnimaa acknowledges financial support from the CHEMS doctoral program of the University of Helsinki.

Declaration of Competing Interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

CRediT authorship contribution statement

Santeri Larnimaa: Investigation, Formal analysis, Writing – original draft, Writing – review & editing, Visualization. **Markku Vainio:** Conceptualization, Writing – review & editing, Supervision, Project administration. **Ville Ulvila:** Conceptualization, Investigation, Writing – review & editing, Supervision, Project administration.

Acknowledgment

We wish to thank Dr. Markus Metsälä and Prof. Lauri Halonen for their support and constructive feedback on the project, and Dr. Guillaume Genoud for fruitful discussions on the topic and for the work to secure funding for the project.

Supplementary materials

Supplementary material associated with this article can be found, in the online version, at [doi:10.1016/j.jqsrt.2021.107984](https://doi.org/10.1016/j.jqsrt.2021.107984).

References

- [1] Simpson WR, Brown SS, Saiz-Lopez A, Thornton JA, von Glasow R. Tropospheric halogen chemistry: sources, cycling, and impacts. *Chem. Rev.* 2015;115(10):4035–62. doi:10.1021/cr5006638.
- [2] Zondlo MA, Hudson PK, Prenni AJ, Tolbert MA. Chemistry and microphysics of polar stratospheric clouds and cirrus clouds. *Annu. Rev. Phys. Chem.* 2000;51(1):473–99. doi:10.1146/annurev.physchem.51.1.473.
- [3] Lique F, Faure A. Collisional excitation and dissociation of HCl by H. *Mon. Notices Royal Astron. Soc.* 2017;472(1):738–43. doi:10.1093/mnras/stx2025.
- [4] Lanza M, Kalugina Y, Wiesenfeld L, Faure A, Lique F. New insights on the HCl abundance in the interstellar medium. *Mon. Notices Royal Astron. Soc.* 2014;443(4):3351–8. doi:10.1093/mnras/stu1371.
- [5] Peng R, Yoshida H, Chamberlain RA, Phillips TG, Lis DC, Gerin M. A comprehensive survey of hydrogen chloride in the galaxy. *Astrophys. J.* 2010;723(1):218–28. doi:10.1088/0004-637x/723/1/218.
- [6] Panu H, Timo R, Thomas F, Makkonen J, Alyshev S, Kharakhordin A, Firstov S. Real-time HCl gas detection at parts-per-billion level concentrations utilizing a diode laser and a bismuth-doped fibre amplifier. *Meas. Sci. Technol.* 2021;32:055206. doi:10.1088/1361-6501/abd651.
- [7] Zuev VV, Zueva NE, Savelieva ES, Gerasimov VV. The Antarctic ozone depletion caused by Erebus volcano gas emissions. *Atmos. Environ.* 2015;122:393–9. doi:10.1016/j.atmosenv.2015.10.005.
- [8] Tamburello G, Caselli A, Tassi F, Vaselli O, Calabrese S, Rouwet D, Capaccioni B, Di Napoli R, Cardellini C, Chiodini G, Bitetto M, Brusca L, Bellomo S, Aiuppa A. Intense magmatic degassing through the lake of Copahue volcano, 2013–2014. *J. Geophys. Res. Solid Earth* 2015;120(9):6071–84. doi:10.1002/2015JB012160.
- [9] Voigt C, Jessberger P, Jurkat T, Kaufmann S, Baumann R, Schlager H, Bobrowski N, Giuffrida G, Salerno G. Evolution of CO₂, SO₂, HCl, and HNO₃ in the volcanic plumes from Etna. *Geophys. Res. Lett.* 2014;41(6):2196–203. doi:10.1002/2013GL058974.
- [10] Pyle DM, Mather TA. Halogens in igneous processes and their fluxes to the atmosphere and oceans from volcanic activity: a review. *Chem. Geol.* 2009;263(1):110–21. doi:10.1016/j.chemgeo.2008.11.013.
- [11] Korabiev O, Olsen KS, Trokhimovskiy A, Lefèvre F, Montmessin F, Fedorova AA, Toplis MJ, Alday J, Belyaev DA, Patrakeev A, Ignatiev NI, Shakun AV, Grigoriev AV, Baggio L, Abdenour I, Lacombe G, Ivanov YS, Aoki S, Thomas IR, Daerden F, Ristic B, Erwin JT, Patel M, Bellucci G, Lopez-Moreno J, Vandaele AC. Transient HCl in the atmosphere of Mars. *Sci. Adv.* 2021;7(7):eabe4386. doi:10.1126/sciadv.abe4386.
- [12] Krasnopolsky VA, Belyaev DA, Gordon IE, Li G, Rothman LS. Observations of D/H ratios in H₂O, HCl, and HF on Venus and new DCl and DF line strengths. *Icarus* 2013;224(1):57–65. doi:10.1016/j.icarus.2013.02.010.
- [13] Sandor BJ, Clancy RT. Observations of HCl altitude dependence and temporal variation in the 70–100 km mesosphere of Venus. *Icarus* 2012;220(2):618–26. doi:10.1016/j.icarus.2012.05.016.
- [14] Rank DH, Rao BS, Wiggins TA. Molecular constants of HCl³⁵. *J. Mol. Spectrosc.* 1965;17(1):122–30. doi:10.1016/0022-2852(65)90114-1.
- [15] Rank DH, Eastman DP, Rao BS, Wiggins TA. Rotational and vibrational constants of the HCl³⁵ and DCl³⁵ molecules. *J. Opt. Soc. Am.* 1962;52(1):1–7. doi:10.1364/JOSA.52.000001.
- [16] Rank DH, Birtley WB, Eastman DP, Rao BS, Wiggins TA. Precise measurements of some infrared bands of hydrogen chloride. *J. Opt. Soc. Am.* 1960;50(12):1275–9. doi:10.1364/JOSA.50.001275.
- [17] Palik ED. History of far-infrared research. I. The Rubens era. *J. Opt. Soc. Am.* 1977;67(7):857–65. doi:10.1364/JOSA.67.000857.
- [18] Ginsburg N. History of far-infrared research. II. The grating era, 1925–1960. *J. Opt. Soc. Am.* 1977;67(7):865–71. doi:10.1364/JOSA.67.000865.
- [19] Imes ES. Measurements on the near infra-red absorption of some diatomic gases. *Astrophys. J.* 1919;50:251–76. doi:10.1086/142504.
- [20] Coxon JA, Hajigeorgiou PG. Improved direct potential fit analyzes for the ground electronic states of the hydrogen halides: HF/DF/TF, HCl/DCl/TCI, HBr/DBr/TBr and HI/DI/II. *J. Quant. Spectrosc. Radiat. Transf.* 2015;151:133–54. doi:10.1016/j.jqsrt.2014.08.028.
- [21] Li G, Gordon IE, Bernath PF, Rothman LS. Direct fit of experimental rovibrational intensities to the dipole moment function: application to HCl. *J. Quant. Spectrosc. Radiat. Transf.* 2011;112(10):1543–50. doi:10.1016/j.jqsrt.2011.03.014.
- [22] Li G, Serdyukov A, Gisi M, Werhahn O, Ebert V. FTIR-based measurements of self-broadening and self-shift coefficients as well as line strength in the first overtone band of HCl at 1.76 μm. *J. Quant. Spectrosc. Radiat. Transf.* 2015;165:76–87. doi:10.1016/j.jqsrt.2015.06.021.
- [23] Asfin RE, Domanskaya AV, Maul C. Broadening and shifting coefficients of rotation-vibrational lines in the fundamental and first overtone bands of HCl and HBr induced by oxygen and air. *J. Quant. Spectrosc. Radiat. Transf.* 2013;130:296–303. doi:10.1016/j.jqsrt.2013.07.014.
- [24] Iwakuni K, Sera H, Abe M, Sasada H. Hyperfine-resolved transition frequency list of fundamental vibration bands of H³⁵Cl and H³⁷Cl. *J. Mol. Spectrosc.* 2014;306:19–25. doi:10.1016/j.jms.2014.09.013.
- [25] Huggle D, Blinov A, Stan-Sion C, Korschinek G, Scheffel C, Massonet S, Zerle L, Beer J, Parrat Y, Gaeggeler H, Hajdas W, Nolte E. Production of cosmogenic ³⁶Cl on atmospheric argon. *Planet. Space Sci.* 1996;44(2):147–51. doi:10.1016/0032-0633(95)00085-2.
- [26] Tosaki Y, Morikawa N, Kazahaya K, Tsukamoto H, Togo YS, Sato T, Takahashi HA, Takahashi M, Inamura A. Deep incursion of seawater into the Hiroshima Granites during the Holocene transgression: evidence from ³⁶Cl age of saline groundwater in the Hiroshima area, Japan. *Geochem. J.* 2017;51(3):263–75. doi:10.2343/geochemj.2.0467.
- [27] Lavastre V, Le Gal La Salle C, Michelot J, Giannesini S, Benedetti L, Lancelot J, Lavielle B, Massault M, Thomas B, Gilibert E, Bourlès D, Clauer N, Agrinier P. Establishing constraints on groundwater ages with ³⁶Cl, ¹⁴C, ³H, and noble gases: a case study in the eastern Paris basin, France. *Appl. Geochem.* 2010;25(1):123–42. doi:10.1016/j.apgeochem.2009.10.006.
- [28] Tosaki Y, Tase N, Massmann G, Nagashima Y, Seki R, Takahashi T, Sasa K, Sueki K, Matsuhiro T, Miura T, Bessho K, Matsumura H, He M. Application of ³⁶Cl as a dating tool for modern groundwater. *Nucl. Instrum. Methods Phys. Res. Sect. B* 2007;259(1):479–85. doi:10.1016/j.nimb.2007.02.096.
- [29] Love AJ, Herczeg AL, Sampson L, Cresswell RG, Fifield LK. Sources of chloride and implications for ³⁶Cl dating of old groundwater, Southwestern Great Artesian Basin, Australia. *Water Resour. Res.* 2000;36(6):1561–74. doi:10.1029/2000WR900019.
- [30] Le Dizès S, Gonze MA. Behavior of ³⁶Cl in agricultural soil-plant systems: a review of transfer processes and modeling approaches. *J. Environ. Radioact.* 2019;196:82–90. doi:10.1016/j.jenvrad.2018.10.011.
- [31] White TL, DiPrete D, DiPrete C, Dobos G. Analysis of ³⁶Cl in Savannah river site radioactive waste. *J. Radioanal. Nucl.* 2013;296(2):835–9. doi:10.1007/s10967-012-2071-9.
- [32] Sheppard SC, Johnson LH, Goodwin BW, Tait JC, Wuschke DM, Davison CC. Chlorine-36 in nuclear waste disposal—1. Assessment results for used fuel with comparison to ¹²⁹I and ¹⁴C. *Waste Manage.* 1996;16(7):607–14. doi:10.1016/S0956-053X(97)00001-9.

- [33] Beasley TM, Elmore D, Kubik PW, Sharma P. Chlorine-36 releases from the Savannah river site nuclear fuel reprocessing facilities. *Groundwater* 1992;30(4):539–48. doi:10.1111/j.1745-6584.1992.tb01530.x.
- [34] Galli I, Bartalini S, Ballerini R, Barucci M, Cancio P, De Pas M, Giusfredi G, Mazzotti D, Akikusa N, De Natale P. Spectroscopic detection of radiocarbon dioxide at parts-per-quadrillion sensitivity. *Optica* 2016;3(4):385–8. doi:10.1364/OPTICA.3.000385.
- [35] Kutschera W. Applications of accelerator mass spectrometry. *Int. J. Mass Spectrom.* 2013;349–350:203–18. doi:10.1016/j.ijms.2013.05.023.
- [36] Hou X. Liquid scintillation counting for determination of radionuclides in environmental and nuclear application. *J. Radioanal. Nucl.* 2018;318(3):1597–628. doi:10.1007/s10967-018-6258-6.
- [37] Fatima M, Hausmaninger T, Tomberg T, Karhu J, Vainio M, Hieta T, Genoud G. Radiocarbon dioxide detection using cantilever-enhanced photoacoustic spectroscopy. *Opt. Lett.* 2021;46(9):2083–6. doi:10.1364/OL.420199.
- [38] Terabayashi R, Saito K, Sonnenschein V, Okuyama Y, Iguchi T, Yamanaka M, Nishizawa N, Yoshida K, Ninomiya S, Tomita H. Mid-infrared cavity ring-down spectroscopy using DFB quantum cascade laser with optical feedback for radiocarbon detection. *Jpn. J. Appl. Phys.* 2020;59(9):092007. doi:10.35848/1347-4065/abb20e.
- [39] Genoud G, Lehmoskoski J, Bell S, Palonen V, Oinonen M, Koskinen-Soivi M, Reinikainen M. Laser spectroscopy for monitoring of radiocarbon in atmospheric samples. *Anal. Chem.* 2019;91(19):12315–20. doi:10.1021/acs.analchem.9b02496.
- [40] Sonnenschein V, Terabayashi R, Tomita H, Kato S, Hayashi N, Takeda S, Jin L, Yamanaka M, Nishizawa N, Sato A, Yoshida K, Iguchi T. A cavity ring-down spectrometer for study of biomedical radiocarbon-labeled samples. *J. Appl. Phys.* 2018;124(3):033101. doi:10.1063/1.5041015.
- [41] Fleisher AJ, Long DA, Liu Q, Gameson L, Hodges JT. Optical measurement of radiocarbon below unity fraction modern by linear absorption spectroscopy. *J. Phys. Chem. Lett.* 2017;8(18):4550–6. doi:10.1021/acs.jpcclett.7b02105.
- [42] McCarth AD, Ognibene TJ, Bench G, Turteltaub KW. Quantifying carbon-14 for biology using cavity ring-down spectroscopy. *Anal. Chem.* 2016;88(17):8714–19. doi:10.1021/acs.analchem.6b02054.
- [43] Genoud G, Vainio M, Phillips H, Dean J, Merimaa M. Radiocarbon dioxide detection based on cavity ring-down spectroscopy and a quantum cascade laser. *Opt. Lett.* 2015;40(7):1342–5. doi:10.1364/OL.40.001342.
- [44] Galli I, Bartalini S, Borri S, Cancio P, Mazzotti D, De Natale P, Giusfredi G. Molecular gas sensing below parts per trillion: radiocarbon-dioxide optical detection. *Phys. Rev. Lett.* 2011;107(27):270802. doi:10.1103/PhysRevLett.107.270802.
- [45] Galli I, Pastor PC, Di Lonardo G, Fusina L, Giusfredi G, Mazzotti D, Tamassia F, De Natale P. The ν_3 band of $^{14}\text{C}^{16}\text{O}_2$ molecule measured by optical-frequency-comb-assisted cavity ring-down spectroscopy. *Mol. Phys.* 2011;109(17–18):2267–72. doi:10.1080/00268976.2011.614284.
- [46] Larnimaa S, Halonen L, Karhu J, Tomberg T, Metsälä M, Genoud G, Hieta T, Bell S, Vainio M. High-resolution analysis of the ν_3 band of radiocarbon methane $^{14}\text{CH}_4$. *Chem. Phys. Lett.* 2020;750:137488. doi:10.1016/j.cpllett.2020.137488.
- [47] Karhu J, Tomberg T, Senna Vieira F, Genoud G, Hänninen V, Vainio M, Metsälä M, Hieta T, Bell S, Halonen L. Broadband photoacoustic spectroscopy of $^{14}\text{CH}_4$ with a high-power mid-infrared optical frequency comb. *Opt. Lett.* 2019;44(5):1142–5. doi:10.1364/OL.44.001142.
- [48] Yim M, Caron F. Life cycle and management of carbon-14 from nuclear power generation. *Prog. Nucl. Energy* 2006;48(1):2–36. doi:10.1016/j.pnucene.2005.04.002.
- [49] Gmitro JJ, Vermeulen T. Vapor-liquid equilibria for aqueous sulfuric acid. *AIChE J.* 1964;10(5):740–6. doi:10.1002/aic.690100531.
- [50] Vol. 83 Griffiths PR, de Haseth JA. *Fourier transform infrared spectrometry. Chemical Analysis: A Series of Monographs on Analytical Chemistry and Its Applications*, 83 John Wiley & Sons; 1986.
- [51] Norton RH, Beer R. New apodizing functions for Fourier spectrometry. *J. Opt. Soc. Am.* 1976;66(3):259–64. doi:10.1364/JOSA.66.000259.
- [52] Savitzky A, Golay MJE. Smoothing and differentiation of data by simplified least squares procedures. *Anal. Chem.* 1964;36(8):1627–39. doi:10.1021/ac60214a047.
- [53] Bernath PF. *Spectra of Atoms and Molecules*. New York: Oxford University Press; 1995.
- [54] Gordon IE, Rothman LS, Hill C, Kochanov RV, Tan Y, Bernath PF, Birk M, Boudon V, Campargue A, Chance KV, Drouin BJ, Flaud JM, Gamache RR, Hodges JT, Jacquemart D, Perevalov VI, Perrin A, Shine KP, Smith M-AH, Tennyson J, Toon GC, Tran H, Tsyuterev VG, Barbe A, Császár AG, Devi VM, Furtenbacher T, Harrison JJ, Hartmann J-M, Jolly A, Johnson TJ, Karman T, Kleiner I, Kyuberis AA, Loos J, Lyulin OM, Massie ST, Mikhailenko SN, Moazzen-Ahmadi N, Müller HSP, Naumenko OV, Nikitin AV, Polyansky OL, Rey M, Rotger M, Sharpe SW, Sung K, Starikova E, Tashkun SA, Auwera JV, Wagner G, Wilzewski J, Wcisło P, Yu S, Zak EJ. The HITRAN2016 molecular spectroscopic database. *J. Quant. Spectrosc. Radiat. Transf.* 2017;203:3–69. doi:10.1016/j.jqsrt.2017.06.038.
- [55] Pine AS, Looney JP. N_2 and air broadening in the fundamental bands of HF and HCl. *J. Mol. Spectrosc.* 1987;122(1):41–55. doi:10.1016/0022-2852(87)90217-7.
- [56] Herzberg G. *Molecular Spectra and Molecular Structure I. Spectra of Diatomic Molecules*. 2nd ed. Princeton: D. Van Nostrand Company, Inc.; 1950.
- [57] Dunham JL. The energy levels of a rotating vibrator. *Phys. Rev.* 1932;41(6):721–31. doi:10.1103/PhysRev.41.721.
- [58] Inostroza N, Letelier JR, Senent ML. On the numerical determination of Dunham's coefficients: an application to $\Sigma^+ \text{X}^1\text{HCl}$ isotopomers. *J. Mol. Struct. Theochem.* 2010;947(1):40–4. doi:10.1016/j.theochem.2010.01.037.
- [59] Rinsland CP, Smith MAH, Goldman A, Devi VM, Benner DC. The fundamental bands of H^{35}Cl and H^{37}Cl : line positions from high-resolution laboratory data. *J. Mol. Spectrosc.* 1993;159(1):274–8. doi:10.1006/jmsp.1993.1124.
- [60] Maslowski P, Lee KF, Johansson AC, Khodabakhsh A, Kowzan G, Rutkowski L, Mills AA, Mohr C, Jiang J, Fermann ME, Foltynowicz A. Surpassing the path-limited resolution of Fourier-transform spectrometry with frequency combs. *Phys. Rev. A* 2016;93(2):021802. doi:10.1103/PhysRevA.93.021802.
- [61] Pine AS, Fried A, Elkins JW. Spectral intensities in the fundamental bands of HF and HCl. *J. Mol. Spectrosc.* 1985;109(1):30–45. doi:10.1016/0022-2852(85)90049-9.
- [62] De Rosa M, Nardini C, Piccolo C, Corsi C, D'Amato F. Pressure broadening and shift of transitions of the first overtone of HCl. *Appl. Phys. B Lasers Opt.* 2001;72(2):245–8. doi:10.1007/s003400000449.
- [63] Tomberg T, Vainio M, Hieta T, Halonen L. Sub-parts-per-trillion level sensitivity in trace gas detection by cantilever-enhanced photo-acoustic spectroscopy. *Sci. Rep.* 2018;8(1):1848. doi:10.1038/s41598-018-20087-9.stylefix.